

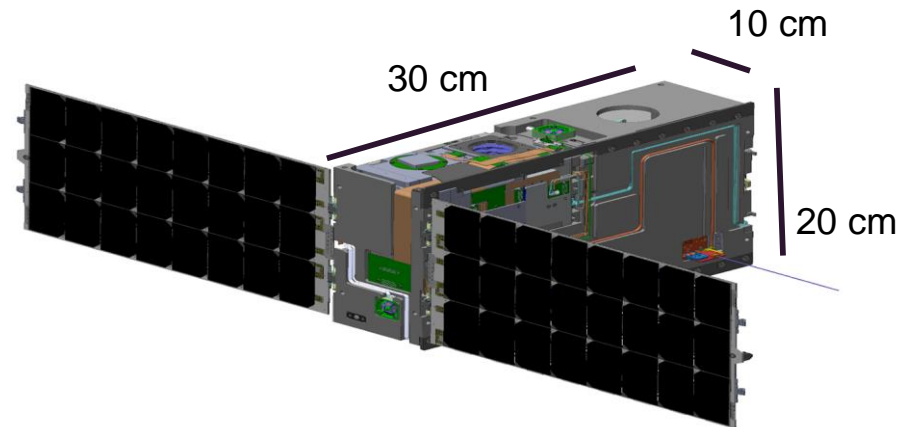
Integration and Testing of the Deformable Mirror (DeMi) CubeSat Payload

Rachel Morgan¹

Gregory W. Allan¹, Ewan Douglas¹, Jennifer N. Gubner^{1, 2}, Yeyuan Xin¹, Paula Do Vale Pereira¹, Bobby G. Holden¹, Christian A. Haughwout¹, Gabor Furesz¹, Mark Egan¹, John Merk³, Kerri L. Cahoy¹

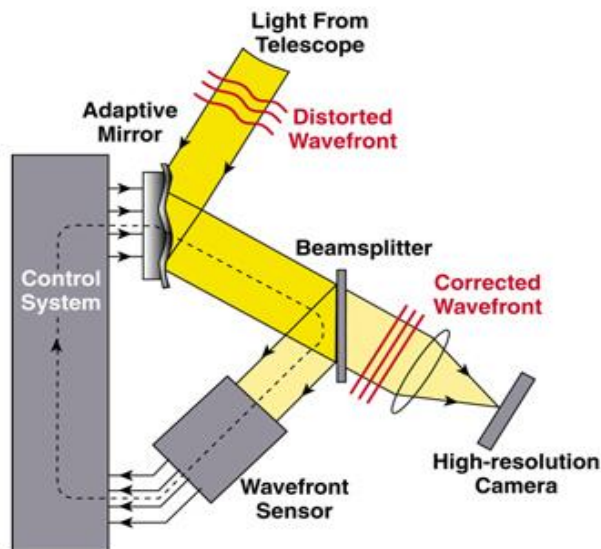
1. Massachusetts Institute of Technology, 2. Wellesley College, 3. Aurora Flight Sciences

- DeMi payload overview
- Integration progress
- Calibration and testing
- Conclusions

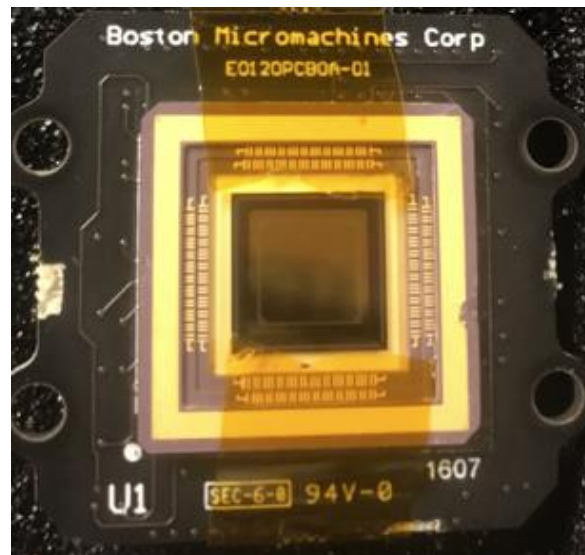


Blue Canyon Technologies

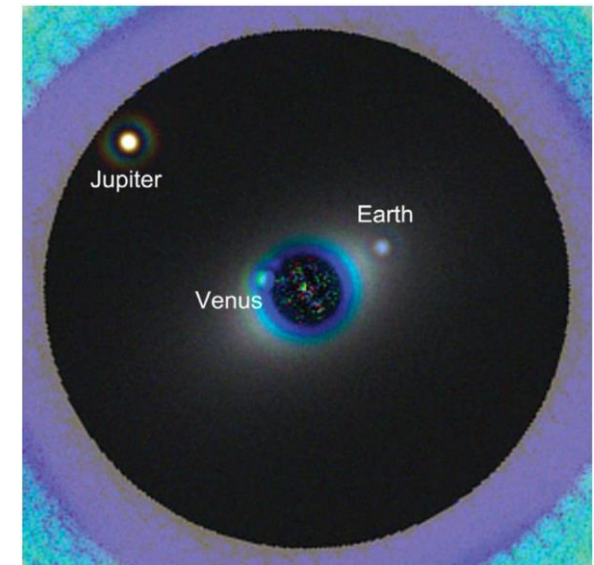
- DeMi: Deformable Mirror Demonstration Mission
- Deformable mirrors correct aberrations and speckles due to mechanical, thermal, and optical effects
- Enables high wavefront stability required for exoplanet direct imaging with coronagraphs



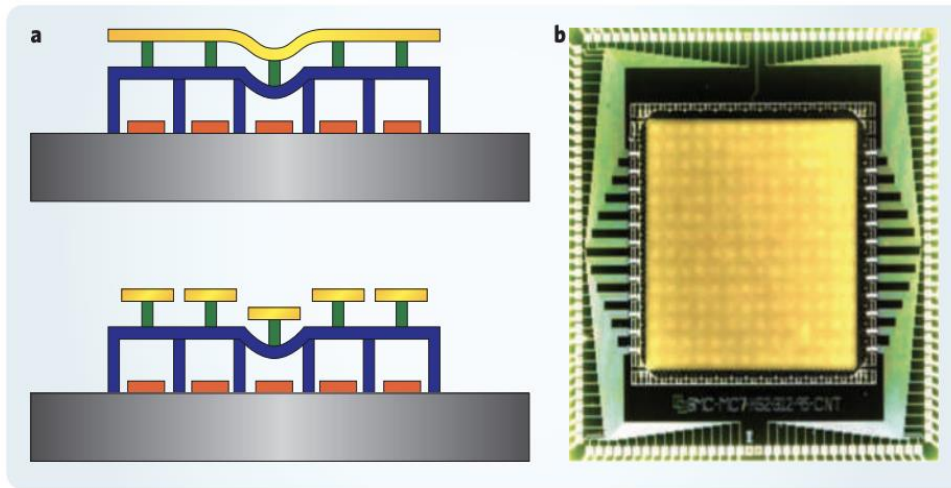
Claire Max CfAO



Boston Micromachines Corporation

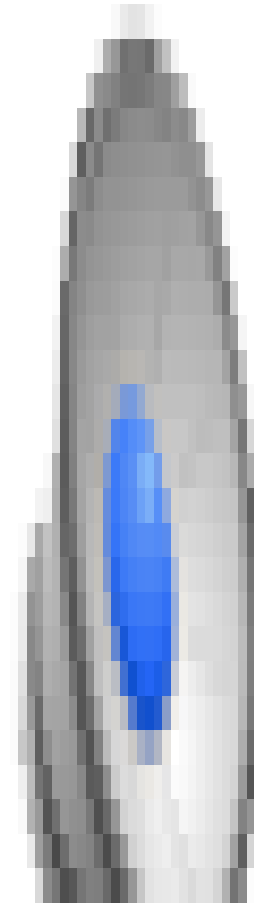


Pueyp and N'Diaye 2015
Simulation of inner solar system imaged through 12m HDST with Apodized Pupil Lyot Coronagraph at 13.5 parsecs



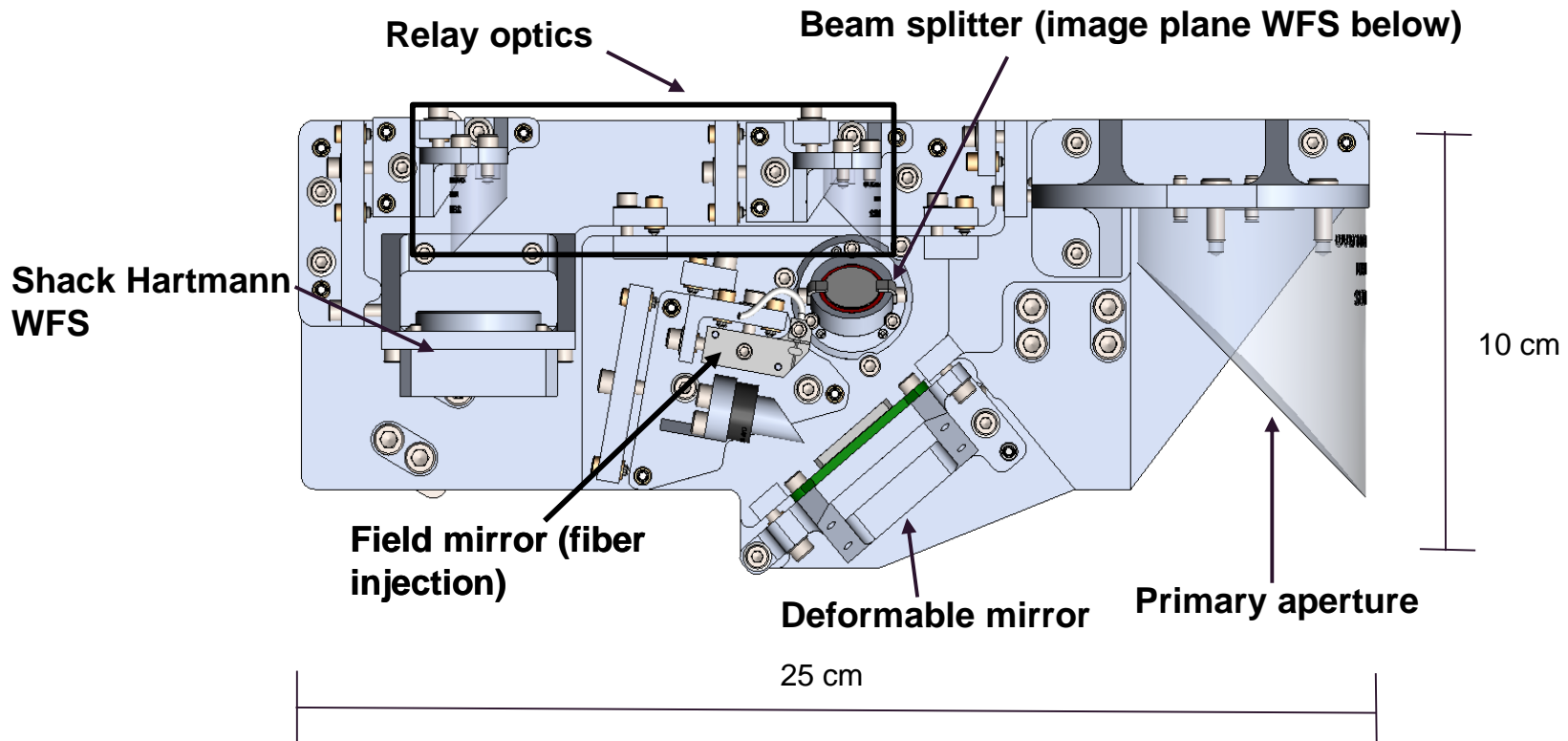
(Above) MEMS DM Design

(Right) Video of poke test of DeMi DM



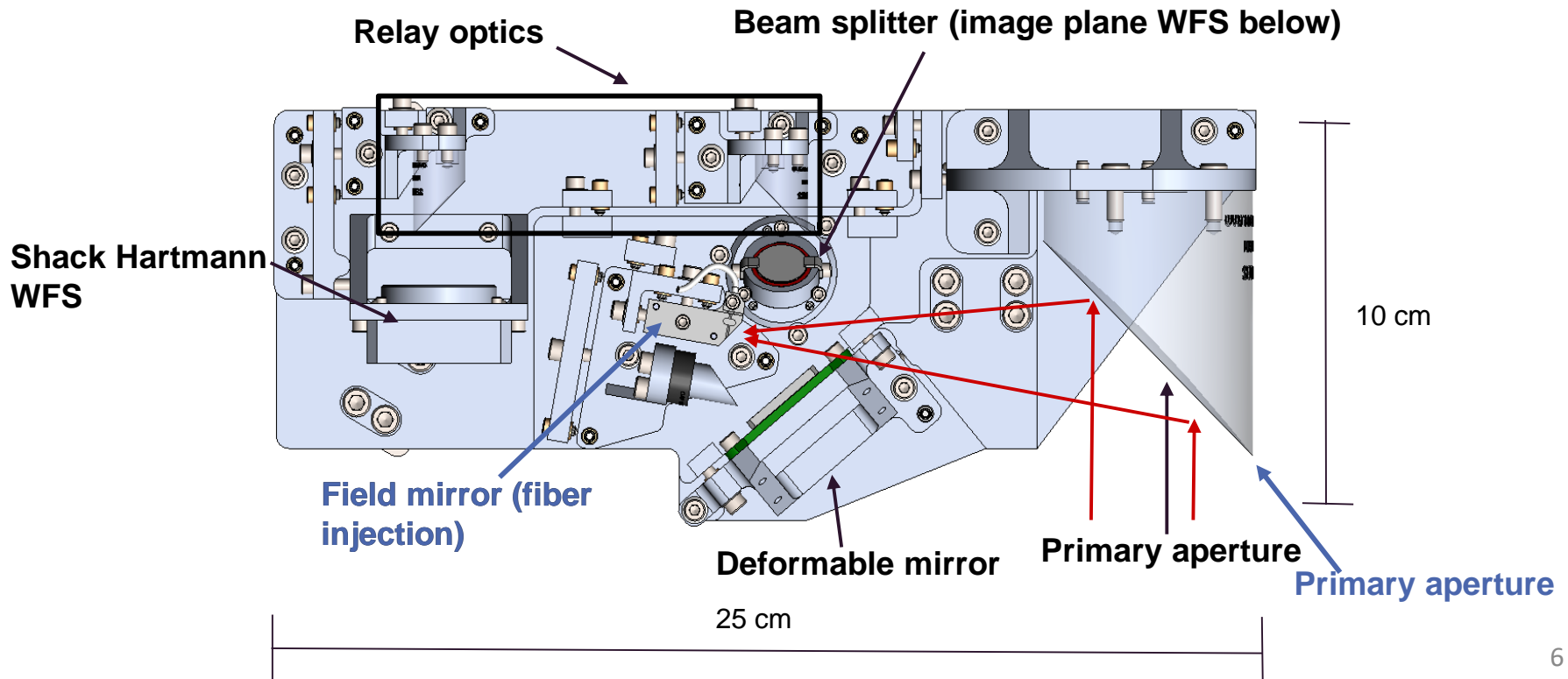
CubeSat mission objectives:

1. On-orbit characterization of 140-actuator BMC MEMS deformable mirror
2. On-orbit demonstration of closed-loop mirror control using both a Shack Hartmann and an image plane wavefront sensor
3. Improve the point spread function (PSF) of an astronomical source



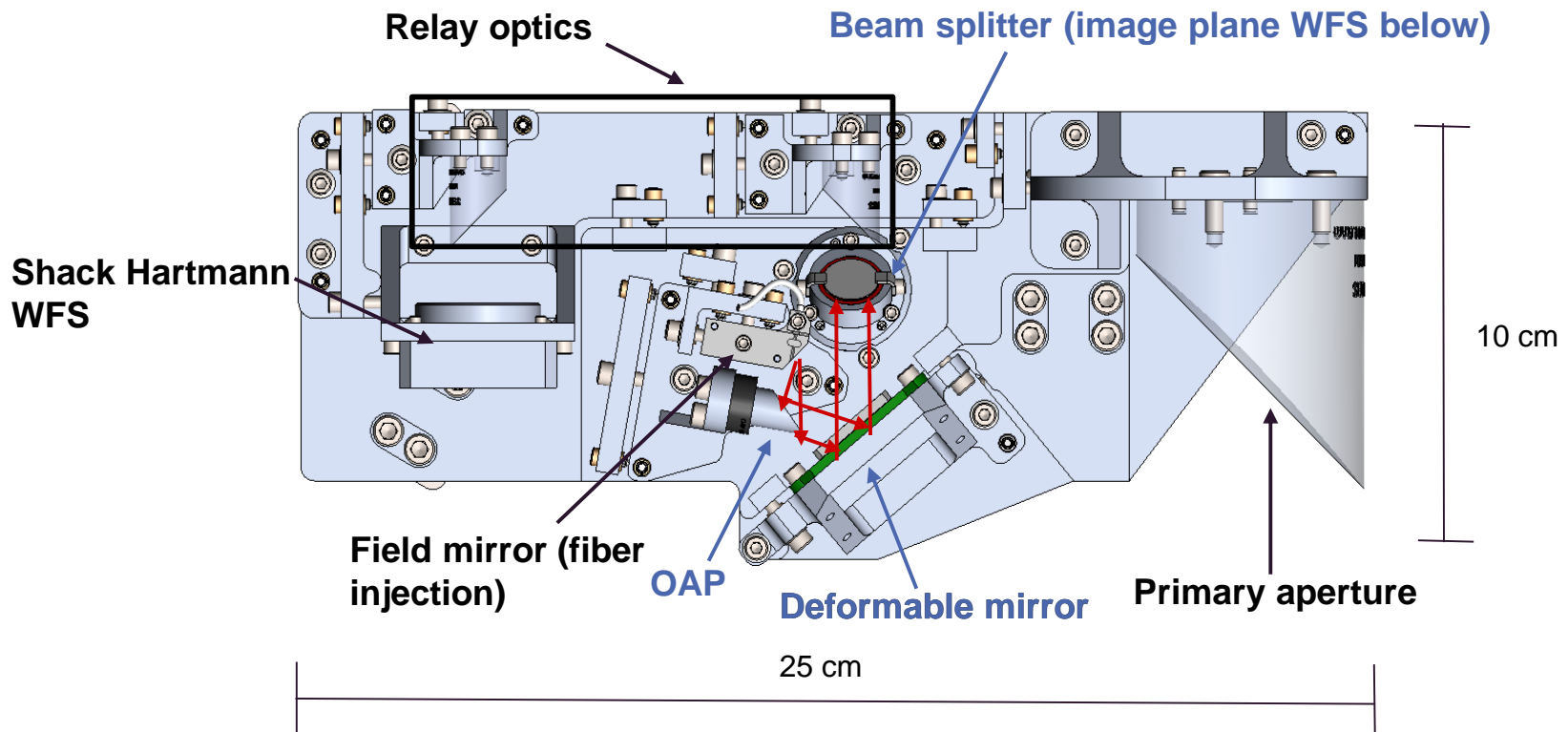
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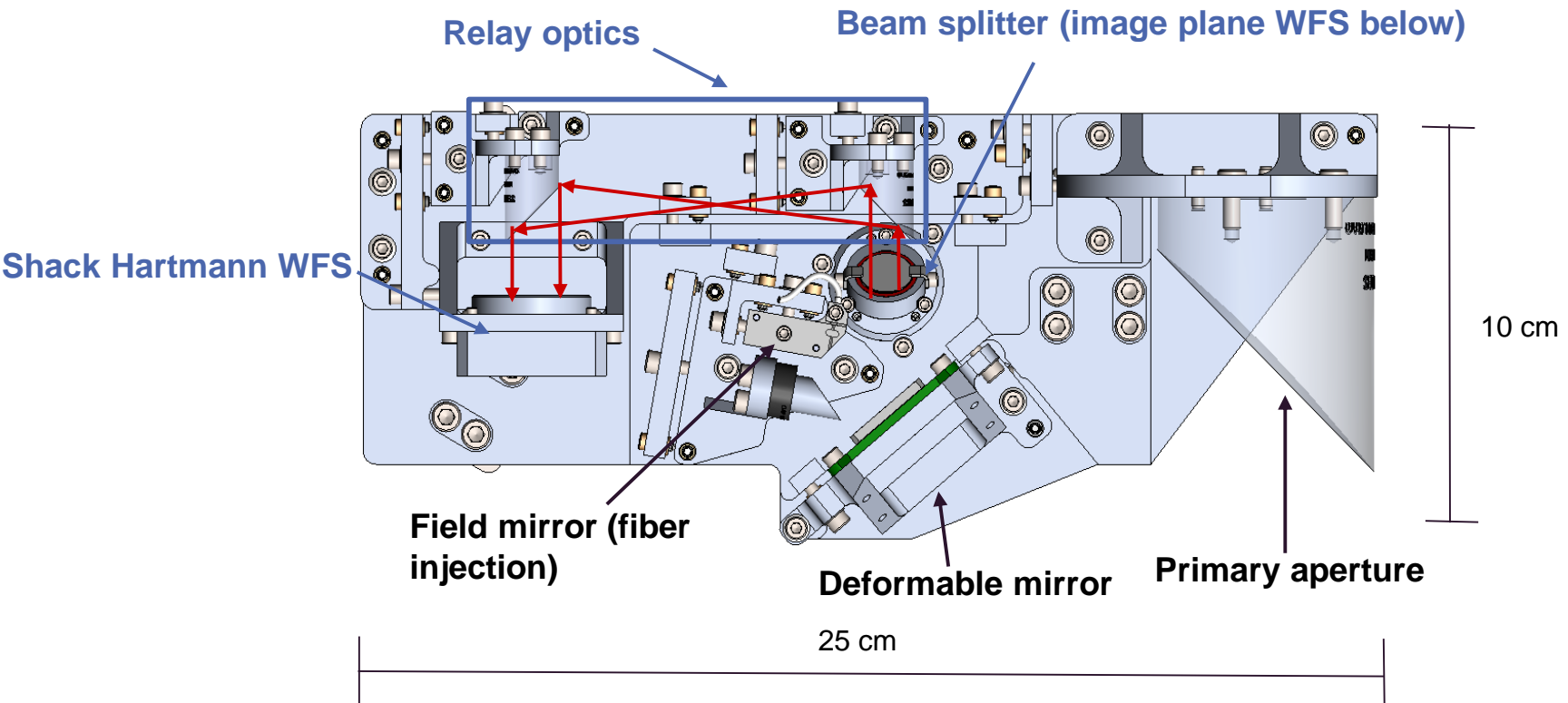
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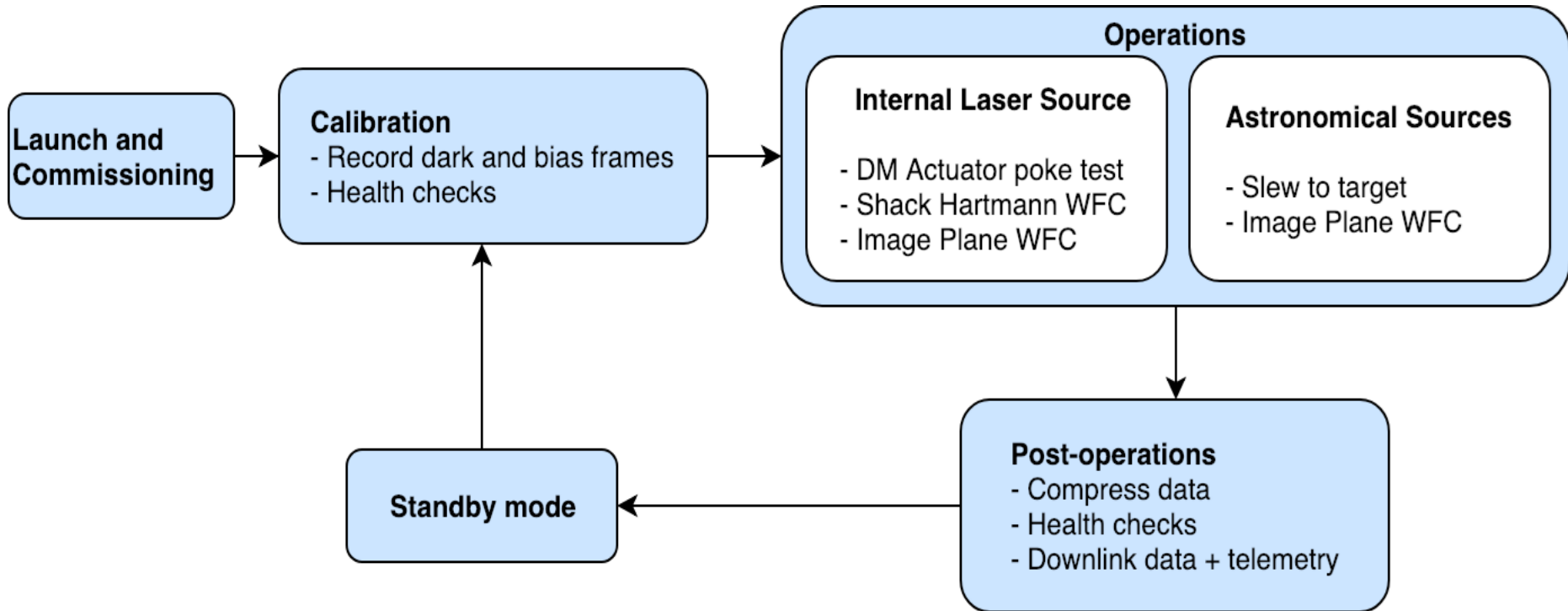
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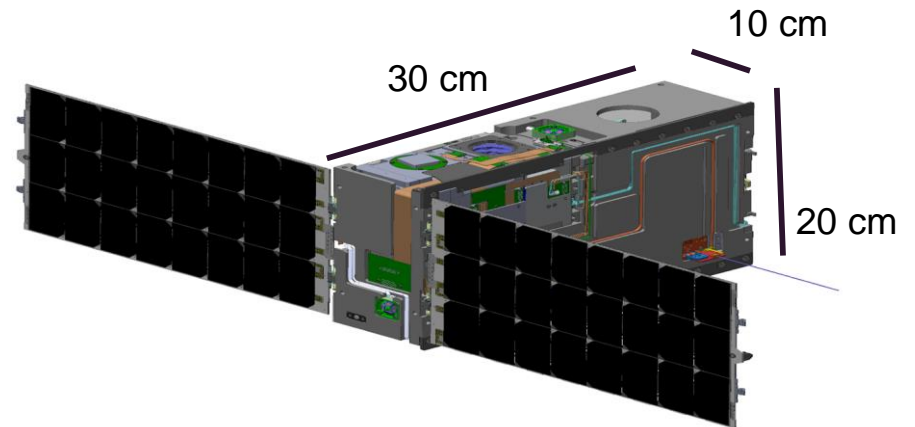


Requirement	Description
L2-SCI-01	The payload shall measure individual deformable mirror actuator wavefront displacement contributions to a precision of 12 nm (TBR).
L2-SCI-02	The payload shall measure low order aberrations to $\lambda/10$ accuracy and $\lambda/50$ precision.
L2-SCI-03	The payload shall demonstrate Deformable Mirror operation for longer than 6 months (TBR).
L2-SCI-04	The payload shall correct static wavefront phase errors up to the DM Nyquist spatial frequency with magnitudes below 3.5 microns to less than 100 nm RMS.
L2-SCI-05	The payload shall correct dynamic wavefront errors up to the DM Nyquist spatial frequency with magnitudes up to 3.5 μm (TBR) to less than 100 nm RMS (TBR).

Requirement	Description
L2-SCI-06	The payload shall measure the PSF of stars brighter than $V_{\text{mag}} = 3$ (TBR) in the image plane with SNR of TBD sufficient to assess the RMS surface error to within 50 nm (TBR) at spatial scales between 0-12 Lambda/D (TBR).
L2-SCI-07	The payload shall be capable of photometric precision of one part in ten thousand (TBR) on stars brighter than 3rd magnitude (TBR) during one orbit.
L2-SCI-08	The payload shall include both image plane and pupil plane sensors to measure the deformable mirror operation in space.
L2-SCI-09	The payload shall be capable of monitoring the current drawn by each high voltage amplifier.



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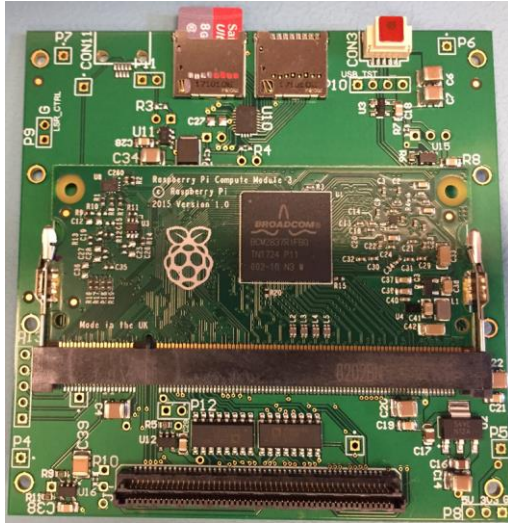
Field mirror
(fiber injection
location)

Beam splitter (Image
plane WFS below)

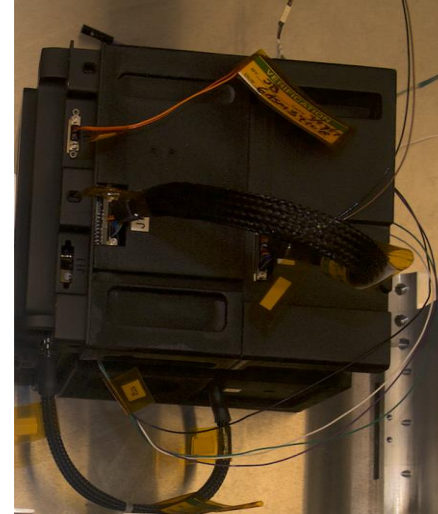
Shack
Hartmann
WFS

Deformable Mirror

Primary
mirror
(OAP1)



2 redundant cross-strapped Raspberry Pi 3s used for payload computers



BCT payload bus simulator for electronics end-to-end testing



23 cm x 18 cm x 6 cm



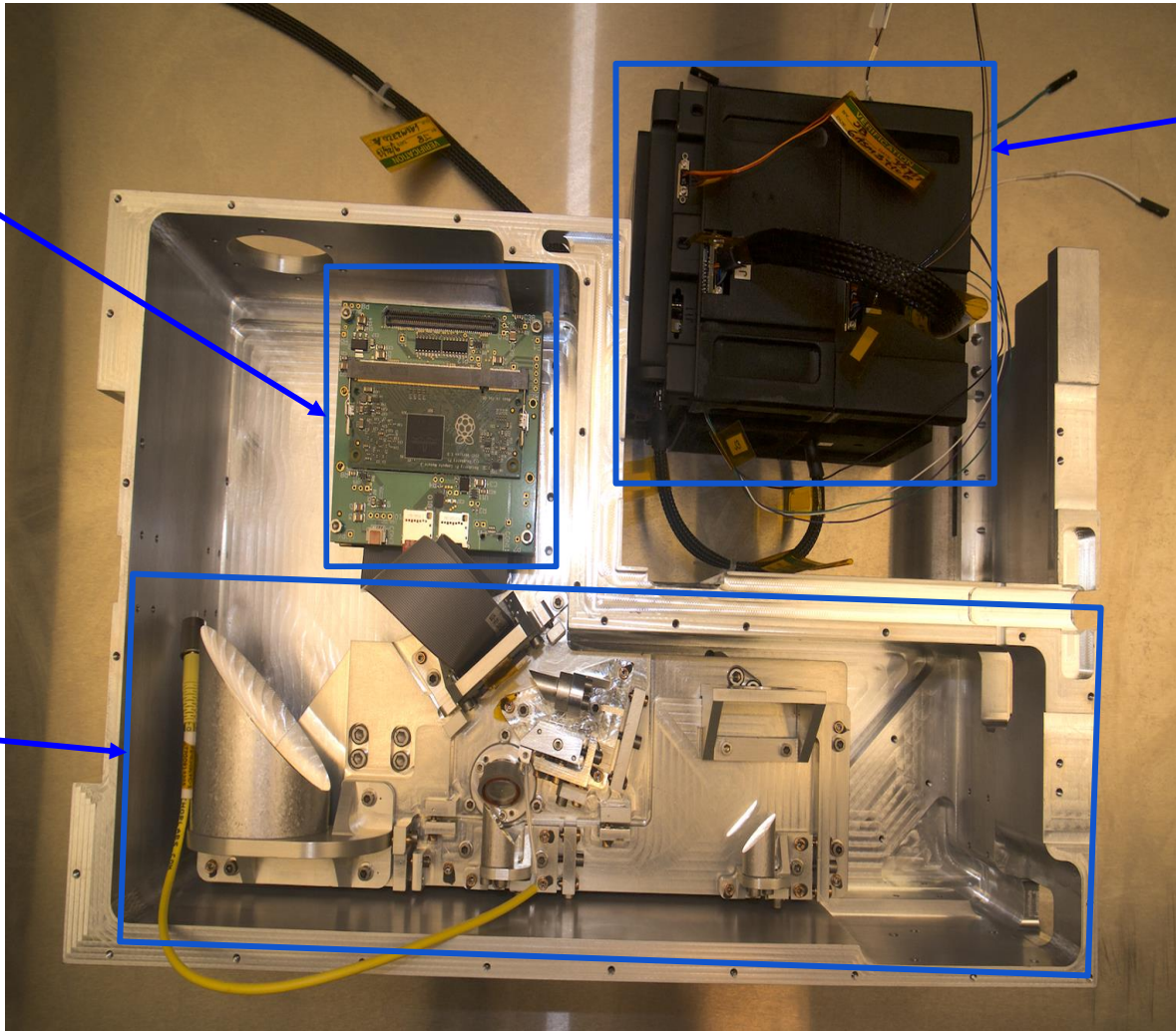
Miniaturized driver electronics to run DM
8 cm x 7 cm x 1 cm

Bus Integration

Payload electronics

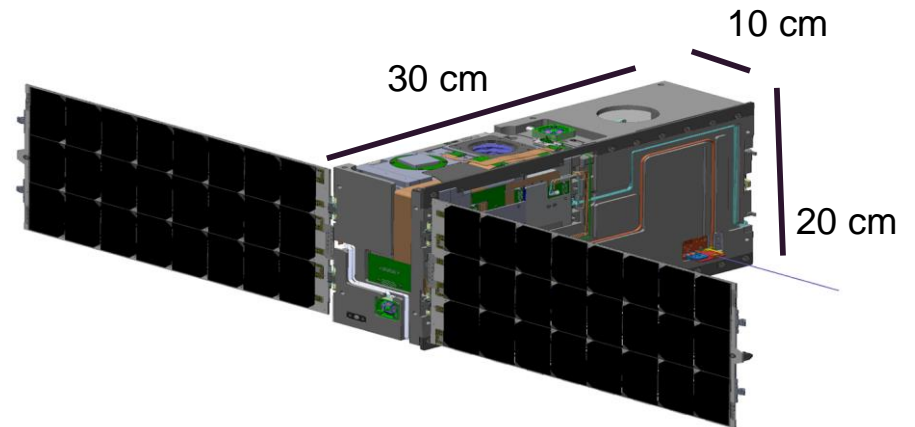
BCT bus simulator

DeMi payload

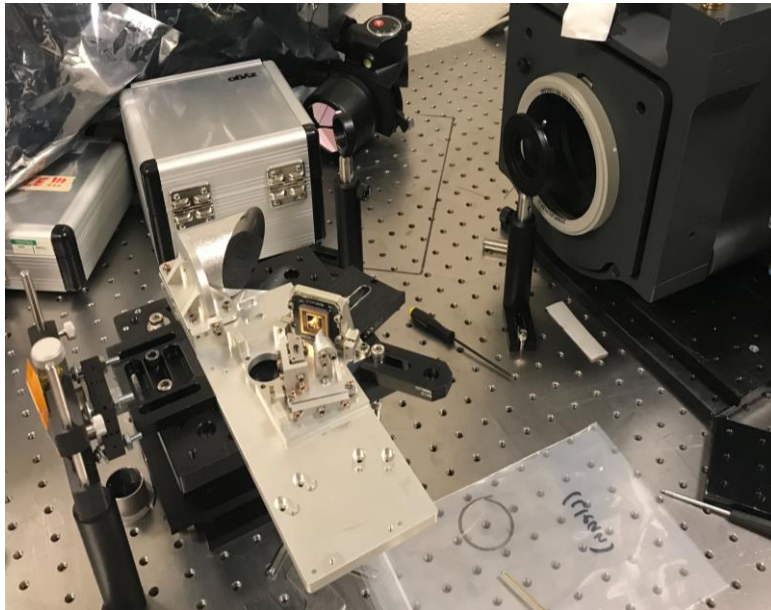


Dimensions: 3U x 6U
Total mass: 12 kg
Payload mass: 2.5 kg

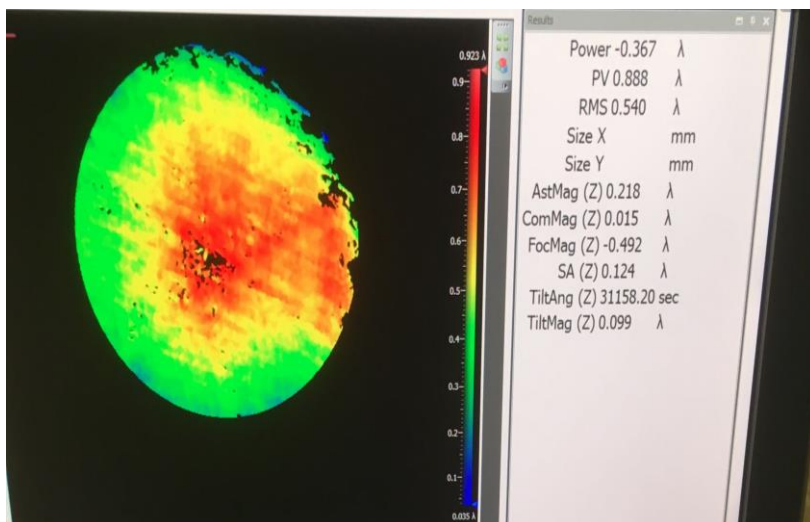
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- Payload aligned with Zygo interferometer
- Started from primary OAP and iterated through system
- Made some mechanical modifications along the way to incorporate into the final flight design



(Top) alignment test setup with Zygo interferometer
(Bottom) example alignment measurement

- Validating wavefront sensing and control with both wavefront sensors using internal laser source
- Results of DM poke tests:

Shack Hartmann WFS

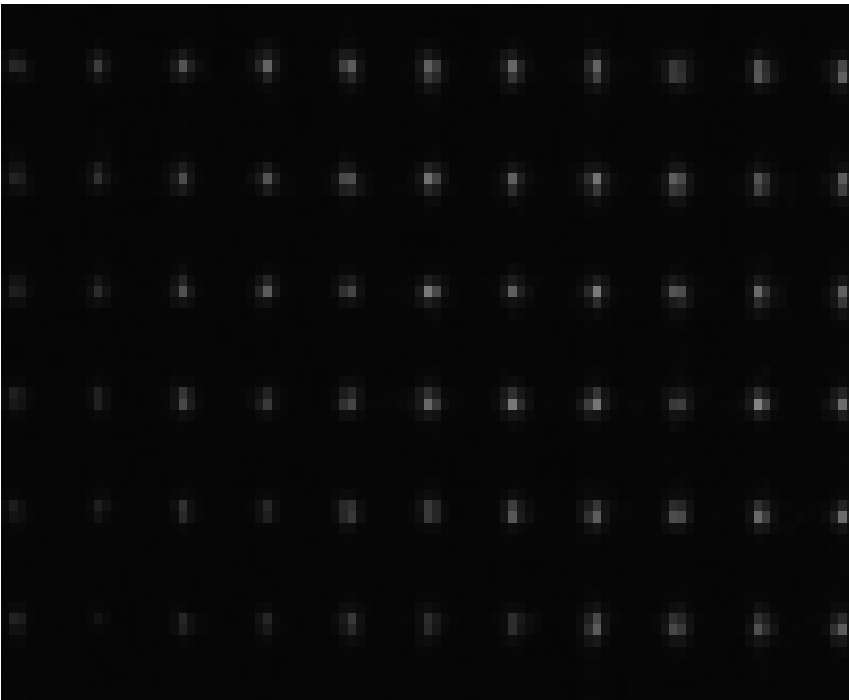


Image Plane WFS



Tip/tilt control of the DM measured on the image plane wavefront sensor



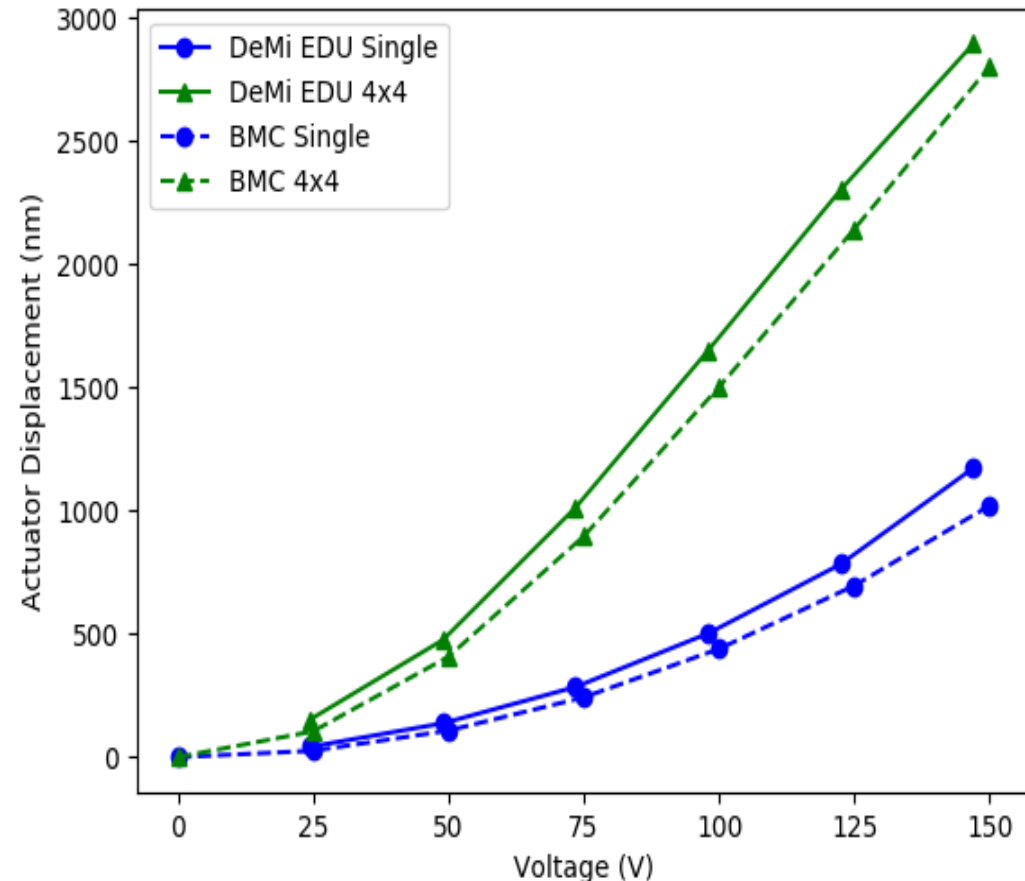
Testing procedures:

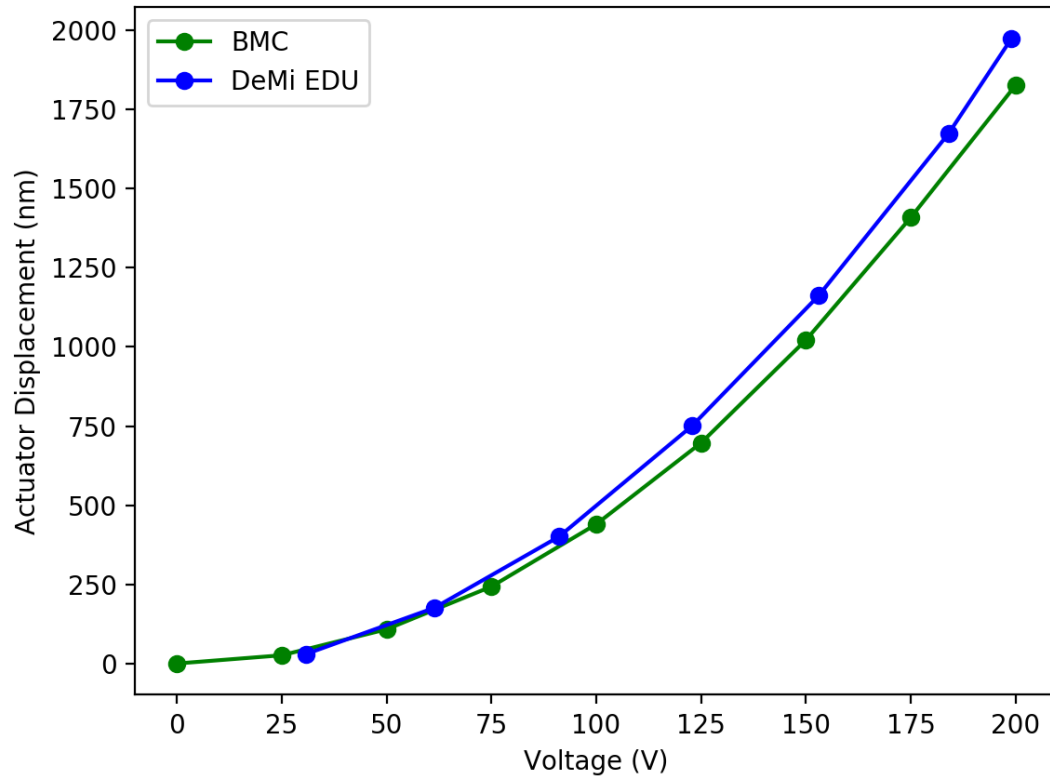
1. Displace **single actuator** (or **4x4 region**) on DM
2. Collect frames on SHWFS
3. Measure spot displacements (average 10 frames, darks subtracted)
4. Plot actuator displacement vs applied voltage, compare to reported values from BMC

Mean differences:

Single actuator – 80.6 nm

4x4 region – 145.1 nm





- Actuator vs displacement curve for single actuator driven by DeMi miniaturized DM driver board
- Compared to BMC reported values
- Mean difference: 67 nm

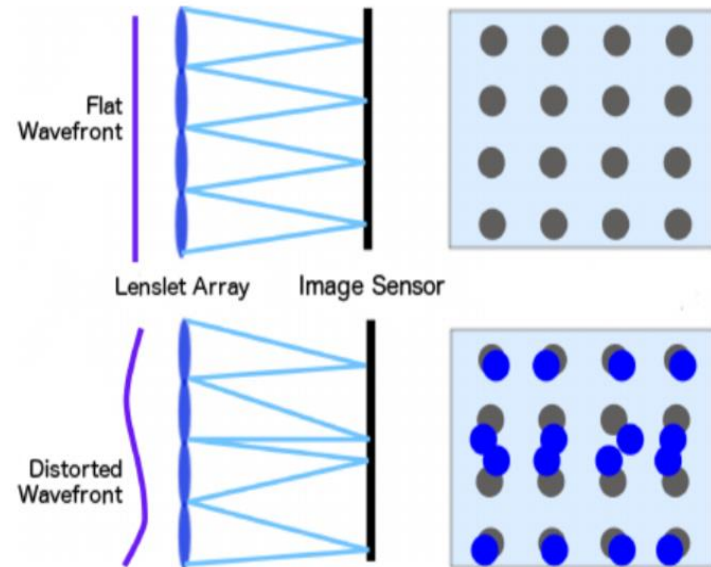
- CubeSat payloads are an excellent platform to test new technologies in space
- DeMi can observe stars through a mini space telescope with AO
- Next steps:
 - More testing of wavefront sensors and DM control
 - Finish flight integration and environmental testing
 - Launch and on-orbit operations!

DeMi is sponsored by DARPA and managed by Aurora Flight Sciences (a Boeing Company)

- Bifano, Thomas. “MEMS Deformable Mirrors.”
Nature Photonics Vol 5(1) 2011.

Parameter	Value
BCT XB6 pointing stability	Sub-10 arcsec 1 sigma pitch and yaw
Deformable Mirror	BMC Multi 140-actuator continuous phase sheet, gold coated
DM stroke	1.5-5.5 um surface displacement
Primary aperture size	30 mm
Telescope magnification	7
Lenslet pitch (SHWFS)	150 um
Effective lenslet focal length (SHWFS)	3.7 um
Mirrors/optical bench material	Aluminum

- Wavefront reconstruction Zonal fit with Southwell geometry
- Singular-value decomposition Moore-Penrose pseudoinverse solution



- Differential Optical Transfer Function
- See Gregory Allan thesis: “Simulation and Testing of Wavefront Reconstruction Algorithms for the Deformable Mirror (DeMi) Cubesat,” MIT 2018

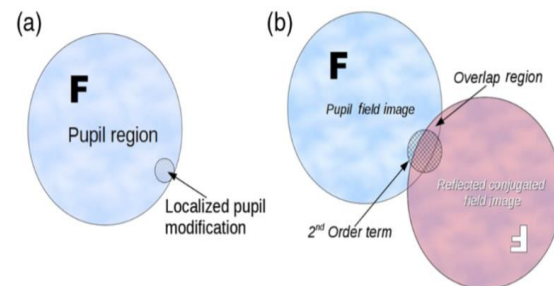


Figure 4-2: Illustration of the spatial distribution of the three terms of the dOTF. [3]

$$\begin{aligned} \delta\mathcal{O}(\xi) &= \mathcal{O}_{\psi+\delta\psi}(\xi) - \mathcal{O}_{\psi}(\xi) \\ &= (\psi + \delta\psi) * (\psi + \delta\psi)^* - \psi * \psi^* \\ &= \psi * \delta\psi^* + \delta\psi * \psi^* + \delta\psi * \delta\psi^* \end{aligned}$$

$$\delta\mathcal{O}(\xi) = \delta\mathcal{O}_L(\xi) + \delta\mathcal{O}_L^*(-\xi) + \delta\mathcal{O}_{\delta\delta}(\xi)$$

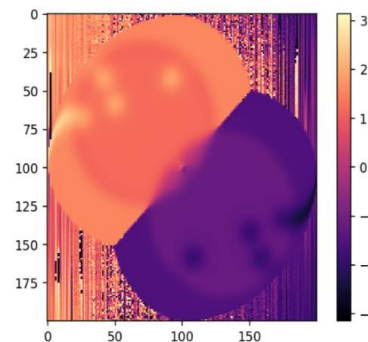


Figure 4-4: Unprocessed, simulated dOTF result, showing the two linear terms of the dOTF are spatially distinct except for overlap at the point of the pupil perturbation. Four actuators are displaced by $\lambda/4$, plus an additional actuator as the dOTF phase perturbation. The expected pupil phase is visible in each term. The spatial axes are given in pixels and the color scale is in radians of wavefront phase.