Simultaneous multi-point space weather measurements using the low cost EDSN CubeSat constellation

Montana State University

Adam Gunderson David Klumpar Matthew Handley Andrew Crawford Keith Mashburn Ehson Mosleh Larry Springer

NASA Ames James Cockrell Hugo Sanchez Harrison Smith





27th Annual AIAA/USU SmallSat Conference August 29, 2013





Introduction



- EDSN Mission
- Instrument Motivations
- Mission Objectives
- Instrument Design
- Instrument Characterization
- Instrument Testing
- Mission ConOps
- Mission Status
- Future Work



Introduction





3



EDSN Mission



- <u>Edison Demonstration of Satellite Networks</u>
 - Multi-node network of 8 identical satellites
 - CMD and DWNLNK accomplished through a single "parent" which sends and receives data from the rest of the "siblings."
 - Any of the 8 satellites can become the "parent."
 - Active attitude control through magnetorquers and reaction wheels
 - On board GPS receiver and S-band Comm system
 - Samsung Galaxy smartphone as CDH







EPISEM Motivations



- 2003 Space Physics Decadal Survey:
 - Multipoint measurements are needed in the ionosphere, where global changes occur on short time scales and small spatial scales.
- 2013 Space Physics Decadal Survey:
 - The study of the heliophysics system requires multipoint observations to develop understanding of the coupling between disparate regions and to resolve temporal and spatial ambiguities that limit scientific understanding.







EPISEM Motivations



EDSN



- EPISEM (above) will measure omnidirectonal integral flux concurrently at each spacecraft
- POLAR/SAMPEX (left) evaluated electron flux at large spatial and temporal distributions
- Spatial variations may occur at scales undetectable by a large spatial distribution.
 - Temporal variations may occur over periods of minutes to hours.

Energetic particles in the Heliosphere and the Magnetosphere, Presentation , Shri Kanekal, LASP



Mission Objectives



How are the spatio-temporal distribution and temporal variability of penetrating electrons and high-energy protons characterized?

- Previous correlations show distributions that appear isotropic in nature when lag time is less than one day, across large spatial/temporal distributions.
- Small spatial/temporal scales accomplished with co-temporal measurements across the EDSN array.



EPISEM will provide the first measurements of coherence at small spatial and temporal scales.



Mission Objectives



What are the fundamental exposure rates of spacecraft avionics to radiation from penetrating electrons and high-energy protons in Low Earth Orbit?

- EPISEM provides constant radiation measurements for each identical spacecraft
- Single event upsets on all or each spacecraft may be correlated to the radiation flux measured by each EPISEM.
- 54% in the South Atlantic Anomaly
- 26% in the Polar Regions
- 20% Galactic Cosmic Rays
- What are the temporal dependencies?
- What are the small scale spatial dependencies?



SEUs in the MOPITT instrument aboard TERRA March 2000 to January 2003



ASEN-5335 Lectures, Jeff Forbes, University of Colorado



Instrument Design



- Employs an thin-walled Geiger-Müeller tube located inside the spacecraft structure.
- Detects penetrating beta/gamma radiation from energetic particles above a certain energy threshold.
- Specific energy threshold is different for electrons and protons.



- Incoming radiation knocks electron off of the Neon fill gas
- Neon becomes Ne+ and free electron avalanches toward anode



- Requires <100 mW of power draw at 8.4V
- Hard-coded serial numbers for all units
- Reports last 60 seconds of counts at 1 second time resolution

August 29, 2013



Characterization



- Plateau curves completed twice for each instrument.
 - After GM tube installed and before shipment (prior to staking / coating).
- High voltage monitor circuit calibrated against a high voltage probe to reduce circuit measurement error.
 - 2nd order curve-fit approximation
- Each board tested using the same procedures.
- Test results compared to ensure no large outliers between boards.



Characterization





Board Serial Number

Board Serial Number



Characterization



EPISEM measures omnidirectional integral flux in counts per second per throughput (GF or $A\Omega$).

- EPISEM: LND71320 Tube, $A\Omega = 1.94$ cm²*sr
- Explorer-1: Anton302 Tube, $A\Omega = 17.4 \text{ cm}^{2*}\text{sr}$
- Omnidirectional flux (ϕ_{part}) = count rate / ($\xi^*A\Omega$)



Integral flux, I(E):

J is a particle flux in counts*sec⁻¹*cm⁻²*sr⁻¹ J has a variation in energy $J(E)dE \propto E^{-\gamma}dE$ EPISEM measures integral flux I(E) above a certain kinetic energy J = -dI/dEAIAA/USU SmallSat Conferen August 29, 2013

EPISEM detects particles with energy greater than some threshold energy.

E > E'

- Found using the CSDA range.
- Approx to the average path length traveled. ٠
- Rate of loss at every point along particles path assumed to be equal to total stopping power.
- Density Cu = 9.0 g/cm^3 , Density Al = 2.7 g/cm^3
- Bottom/Side Entry -> 2 MeV e- and 25 MeV p+

Top of Stools Entry	Material	E' for e-	E' for p+
TOP OF STACK ENTRY	wateria	(MeV)	(MeV)
S-Band Patch	Copper	2.00	25.00
Router Assembly PCB	Copper	2.00	25.00
S-Band Heatsink	Aluminum	4.50	13.00
MHX2420 Shield	Aluminum	4.50	13.00
MHX2420 PCB	Copper	2.00	25.00
ACS PCB	Copper	2.00	25.00
Phone PCB	Copper	2.00	25.00
EPS PCB	Copper	2.00	25.00
18650 Liions	Various	8.00	60.00
Stensat Radio PCB	Copper	2.00	25.00
GPS Aluminum Housing	Aluminum	4.50	13.00
GPS Card	Copper	2.00	25.00
GPS Heatsink	Aluminum	4.50	13.00
GPS Interface PCB	Copper	2.00	25.00
EPISEM PCB	Copper	2.00	25.00
Geiger Tube Wall	Aluminum	0.08	2.20
Total Thicknesses	Copper	20.00	95.00
Total Thicknesses	Aluminum	1.75	27.50 2



Testing





- TVAC chamber test:
 - Hot soak at 50C
 - -40C to 40C cycles
- High Altitude balloon flight:
 - Flight had two other detectors
 - Maximum altitude of 28 km

Suspected heavy ion strike from measurements with a silicon detector set to a threshold of >40 MeV



Geiger Counter Data



Sterre & Depending Lab





Science Data Collection:

Current baseline is 30 min data runs, once per day.

On-board propagator used for location targeting



S-band Crosslink/Downlink:

"Siblings" crosslink to "Parent" "Parent" downlinks during pass Santa Clara Ground Station August 29, 2013



HAM community packets for E1P/HRBE:

One packet = 150 bytes

25,000 HAM packets received total.

Points represent number of packets received when the satellite is at that particular location.



EDSN's round robin-beacon structure.

Signed Science & Explorering Lab

ConOps



• Targets of interest

- L-shell conjunctions: RBSP, BARREL
- Flares, CME's, solar proton events
- South Atlantic Anomaly passes

Flare initially blamed for the Galaxy IV satellite failure



²⁵ Oct 2013 07:51:34.064 Time Step: 0.001

August 29, 2013

AIAA/USU SmallSat Conference



Data Assimilation in the Radiation Belts, D.N. Baker, E.J. Rigler, et al., 2003, LASP









- All instruments delivered to Ames
 - QA inspection passed
 - 8µC, Cs-137 radiation source procured at Ames
 - Awaiting pre and post integration performance tests
- Launch mid-2014 on ELaNA VII

Questions?

Acknowledgements

NASA Ames Research Center Andrew Crawford, Matthew Handley, Jerry Johnson, David Klumpar, Larry Springer, Keith Mashburn, and Ehson Mosleh







17

April 19, 2012





Backup Slides

Instrument Design















August 29, 2013

AIAA/USU SmallSat Conference







- All circuit boards cleaned, cleaned, and then cleaned again
- Boards staked and coated before final testing and shipment







- Mounted and packaged in an aluminum shipping container
- **Double-wrapped in alumaloy**
- Packaged and shipped in lots of two to four







EPISEM BOREALIS Flight Counts per minute moving 10 second average



Design Iterations



- 0 ×











VBATT Amperage (mA)











Temperature 1



AIAA/USU SmallSat Conference

Design Iterations





August 29, 2013

AIAA/USU SmallSat Conference



Space Weather



Customer Growth SWPC Product Subscription Service



Customers — Solar Cycle