



A Stellar Gyroscope for CubeSat Attitude Determination

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Presentation Overview

- Stellar Gyroscope Concept
- Motivation
 - CubeSat ADCS
 - Attitude knowledge in eclipse
- How the Stellar Gyroscope Works
- Flight Experiment

Concept of Stellar Gyroscope



Observe the motion of stars in camera's field of view to infer changes in satellite's attitude.



- Measures relative attitude between exposures with common stars
- Tolerates large amount of noise, allowing low cost assembly and small form factor

CubeSat Attitude Determination Challenge

In Sun Light
Magnetic Field Vector
Sun vector
Earth Sensor
Star Tracker: needs baffle

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<u>In Eclipse</u>

Magnetic Field Vector
 No Sun Vector

Earth not lit for Earth
 Sensor

🔹 Star Tracker

Rate Gyroscope integration with drift

SSBV CubeSat Solution: Sun: Sun Sensors & Magnetometer + GPS Eclipse: MEMS Gyros + **Stellar Gyro**

Motivation - Alternatives for Eclipse

- Attitude measurement challenging (Earth sensor in InfraRed, Star Tracker)
- Laser Ring Gyros: Highly accurate, Large, Expensive

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- MEMS Gyroscopes:
 - Compact, and Affordable
 - Small Satellites almost exclusively use MEMS
 - Noisy: drift ~0.5 degrees per minute
- Image-based Approach (stellar gyro):
 - Comparable volume and cost
 - Added computational requirements
 - Can assist MEMS gyros by limiting drift





5/18





ADCS System

- In Sun light:
 - Magnetometer, with magnetic model and GPS position knowledge
 - Sun Sensors, using Position Sensitive Detector (not photodiodes)
- In Eclipse:
 - MEMS inertial gyroscopes
 - Assisted by stellar gyroscope to reset drift



Inertial Gyros

Camera Specifications

- CMOS Sensor with S-Mount Lens
- Designed to capture Star Magnitude 4 and brighter
- At least 3 stars in Field of View in 99% of the sky.

Parameter	Value
Sensor	OmniVision OV7725 CMOSVGA Sensor (640 x 480 pixels)
Optics	6 mm focal length, Aperture F/2.0
Field of View	27.6° by 36.7°
Sensitivity	3.8 V/(Lux · s)
S/N Ratio	50 dB
Dark Current	40 mV/s
Pixel Size	6 x 6 µm





Camera assembly as experiment on TDS-1. Further miniaturization is possible.



Star Detection

"Centroiding", aka Expected Value





Camera Model

- Camera Calibration to characterize:
 - Principal Point
 - Focal Length
 - Lens Radial and Tangential Distortion
- Used to acquire precise star vectors





Camera Calibration Toolbox – by Jean-Yves Bouguet http://www.vision.caltech.edu/bouguetj/calib_doc/

400

200

9/18

200

0

-100

-200

Using the Direction-Cosine-Matrix (DCM) notation, the attitude change between two frames satisfies:

$$\overrightarrow{\mathbf{v}^{b}} = C^{ba} \overrightarrow{\mathbf{v}^{a}}$$

- The goal is to find the rotation matrix (C^{ba}) that defines the rotation between frame a and frame b.
- Given at least two vector measurements (two stars before-and-after), The Q-Method is used to find the analytically optimal relative attitude estimate.



Correspondence Across Frames

- False-positives: noise
- False-negatives: missed stars
- Entering and Leaving FOV.
- Correspondence Problem: identifying the same star across frames.
 - By brightness: highly susceptible to noise
 - By predicted location: susceptible to unexpected maneuvers and to false-stars and missed stars



Overlaid detected stars in 5 images that are 3 degrees apart

Random Sample Consensus (RANSAC)



- RANSAC: iterative method to estimate parameters of a mathematical model from a set of observed data which is contaminated a large number of outliers that do not fit the model.
- The steps of RANSAC can be summarized as
 - Hypothesize: A hypothesis rotation is based on MEMS rate information, or calculated using randomly selected star pairs across frames.
 - **Test:** The estimated rotation matrix is tested against all the stars in the two frames. Stars that show consensus are counted towards the Consensus Set (CS).
 - Iterate: RANSAC iterates between the above two steps until a random hypothesis finds "enough" consensus to some selected threshold.

RANSAC Performance

- Observe Motion of Earth
- I degree every 4 minutes
- Photos of the night sky at 0.25° increments, pointing arbitrarily up.
- Prototype Hardware

Rotation Estimate = 1.4495° , Actual Rotation = 1.5°



 Photos of the night sky using Spin Table

Angle Estimate = 25.0549°, Actual = 24.960975°

Canon GI0



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Attitude Response in Eclipse: MEMS only



- Assuming perfect attitude knowledge before entering eclipse
- MEMS rate gyro: 50Hz, ±80 °/second, 12-bit ADC, Noise 0.1
 °/second RMS
- Attitude knowledge error increases up to 5° in the first
 5 minutes and more than 10° after 35 minutes.





MEMS assisted by Stellar Gyroscope

- Assuming perfect attitude knowledge before entering eclipse
- Stellar gyro generates attitude estimates (σ = 0.1°), at 15 second increments, relative to the first photo taken at the beginning of eclipse.
- Drift is maintained below 1°





TechDemoSat-I

- Surrey Satellite Technology LTD, UK. Around 1-meter cubed, 150 kg.
- No less than 8 technology demonstration payloads Maritime Suite, Space Environment Suite, Air and Land Monitoring Suite, Platform Technology Suite
- TDS-I will test CubeSat ACS payload developed by SSBV Space and Ground Systems UK
- KySat-2 (Kentucky Satellite-2)
 - Kentucky Space Consortium
 - I-Unit CubeSat
 - Improved refight of KySat-I mission objectives



- Stellar Gyroscope finds relative attitude by tracking stars
- Correspondence (using RANSAC) can be done with large levels of noise, enabling implementation with low cost sensors and optics
- SSBV CubeSat ADCS system is designed to maintain high quality attitude knowledge throughout the orbit
- In Eclipse, it uses a Stellar Gyroscope to reset the drift of a MEMS attitude propagator.

Thank You

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