COSMIC X-RAY BACKGROUND NANOSAT

Spring 2011 CubeSat Developer's Workshop

Tyler G. Rose, Student Project Lead Benjamin K. Malphrus, Mentor Kevin Z. Brown, Mentor



MISSION – TEAM

• Morehead State University (MSU)

Benjamin Malphrus, Kevin Brown, Bob Twiggs, Jeff Kruth, Mike Combs, Robert Kroll, Thomas Pannuti, Margaret Powell, Andrew Cavins, Jedidiah Reader, Brad Schneider, Tyler Rose

UC Berkeley

Garrett Jernigan

• Noqsi Aerospace, Ltd.

John Doty, Matthew Wampler-Doty

Lawrence Livermore National Laboratories (LLNL)

Lance Simms

• Sonoma State University (SSU)

Steve Anderson, Lynn Cominsky, Kamal Prasad



MISSION – DESCRIPTION

- Very few flux measurements of high energy cosmic X-ray background have been made
- Goal: Make precise measurements of the cosmic X-ray background (CXRB)
 - Intended range: 30-50 keV
 - Valuable for many reasons:
 - Ensure proper background subtraction when studying objects near the same emission (outskirts of clusters and groups of galaxies)
 - About 30% of the emission is unaccounted for, even considering discrete sources and clustered phenomena
 - Origin of missing flux?
 - Lends insight into the evolution of primordial galaxies



MISSION – X-RAY BACKGROUND

Slides Courtesy of Lance Simms, LLNL

NGC 4151 A bright Active Galactic Nucleus (AGN) outside our galaxy



Position

CXBN - PDR

Flux

MISSION – X-RAY BACKGROUND

We know where this intensity is coming from

But what about this Diffuse Background?

4/21/2011

Position

CXBN - PDR

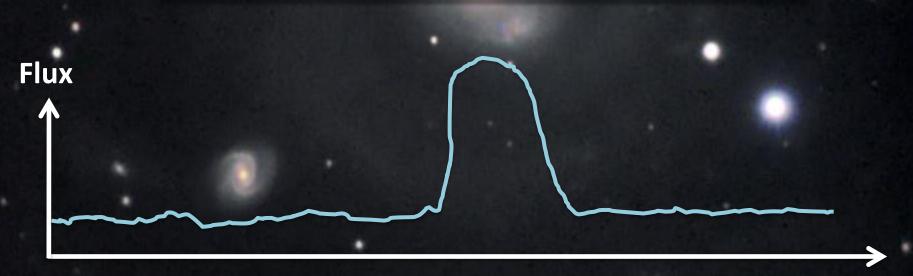
Flux

MISSION – X-RAY BACKGROUND

Is the Cosmic X-Ray Background (CXRB) due to AGN that are too distant to resolve with our telescopes or something else?

"... the cosmic X-ray background (CXRB), still remains one of the most interesting topic of X-ray astronomy and observational cosmology."

- Revnivtsev et al., A&A 444, 2005



Position



MISSION – DESCRIPTION

- Selected for ELaNa VI with expected flight for 2012
- Orbit
 - Moderately-high inclination
 - Altitude approximately 500 km
- 2U CubeSat
 - 1U Payload
 - 1U S/C
- Sun-pointing



MISSION – CONSTRAINTS

- Time project completion date scheduled for Dec. 2012
 - Keep it simple
- Drive costs to a minimum
 - Keep it simple
 - Maximize subsystem reuse
 - Utilize internal resources prototyping, manufacturing, testing
- 3 degrees, 3 sigma accuracy for attitude knowledge
 - Post-process ADCS sensor data to reduce payload data
 - Spin stabilize



PAYLOAD

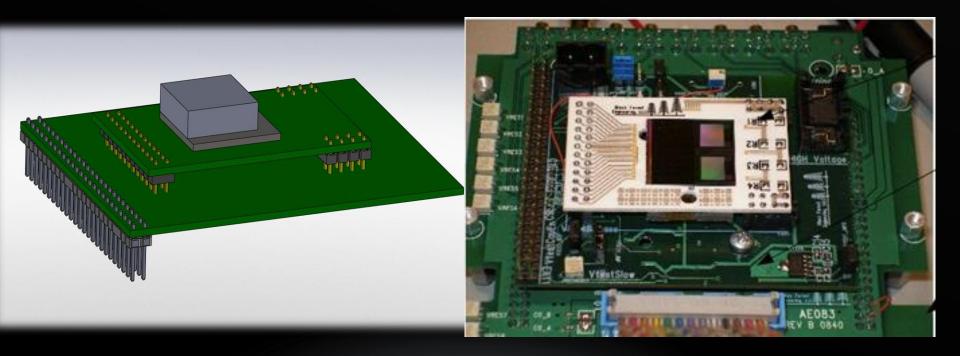
- Cadmium Zinc Tellerium (CZT) Detector Array
 - Goal: < 5% flux measurement precision
 - Designed by UC Berkeley, LLNL, Noqsi Aerospace, and SSU
 - Features a collimator for preventing stray particles
 - Shielding layered lead, copper, and tantalum



PAYLOAD

Early CAD model of payload

CZT Array on development boards





ADCS

- Determination
 - Sun Sensors
 - Coarse (CSS)
 - Medium (MSS)
 - Fine (FSS)
 - Canopus Pipper
 - Magnetometer
 - MEMS Gyro

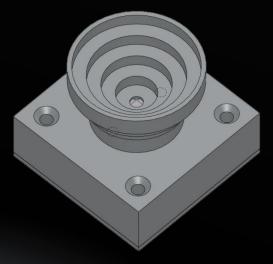
- Control
 - Magnetorquers
 - Simulink simulation for control tuning



ADCS – DETERMINATION

- Sun Sensors
 - Coarse, Medium, and Fine
 angular resolution
 - Sun pointing
 - Boresight
 - Testing through rapidprototyped parts
- (Right) CAD Model
- (Right Top) Rapid-prototyped sun sensor assembled

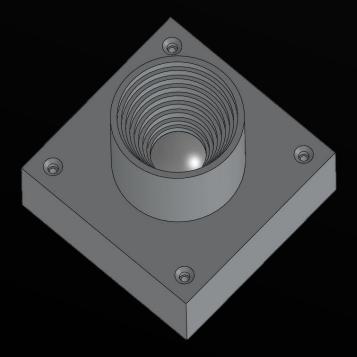






ADCS – DETERMINATION

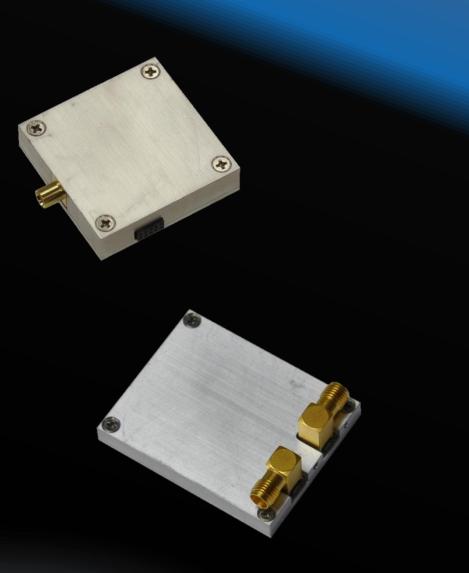
- Canopus Pipper
 - Star Sensor
 - Brightest star in S.Hemisphere
 - Determine roll rate
 - Testing
 - Rapid-prototype frame and collimator
 - Calibrate on ground
- (Right) CAD Model





COMM

- Radios:
 - AstroDev CII:
 - UHF-UHF Transceiver
 - AstroDev Beryllium 2:
 - S-Band Transmitter
- Antennas:
 - Steel tape whip antennas





C&DH / PROCESSOR

- Texas Instruments:
 - OMAP 3530
 - ARM Cortex-A8, 720 MHz
 - Linux OS
 - Power efficient



POWER

- Body mount solar cells and fold out panels
 - Sun pointing: moderate incident angles no MPPT
- Batteries lithium ion
 - Higher capacity
- In-house designed direct energy transfer
 - Keep it simple
 - Quick turn-around



STRUCTURE

- Complete CAD model before manufacturing
 - Leverages previous design work
- Rapid-prototyped and machined onsite
 - Fit checks
 - Quick turn-around
 - Lower costs

GROUND STATIONS

- 21 meter Space Tracking Antenna at MSU
 - S-Band communication
- Dual Band Yagi Antenna
 - VHF/UHF communication
- Portable Satellite Tracking Kits
 - Primarily an E / PO effort
 - Dual Band Yagi Antennas





ESTIMATED TIMELINE

Cosmic X-Ray Background Nanosat (CXBN) Mission Timeline												
		2011										
#	Event	Jan	Feb	Mar	Apr	May	Jun	Jul	Aug S	ep Oct	Nov	Dec
1	Science Mission Concept											
2	Spacecraft Design Concept											
3	Payload Requirements Defined											
4	Design / Prototype Phase											
5	Development / Production Phase											
6	Preliminary Design Review											
7	Eng. Model Prototype & Production											
8	Eng. Model Functional Testing											
9	Interface Control Document Received											
10	Critical Design Review											
11	Flight Unit Orders											
12	Flight Systems Fabrication / Assembly											
13	Flight Payload Fabrication / Assembly											
14	Flight Payload Integration											
15	Flight Model Functional Testing											
16	Test Results Provided			1								
17	Preflight Testing onsite (2U Only)											
18	Preflight Testing offsite (2U Only)											
19	Comprehensive Test Report											



CONCLUSION

- CXBN team is very excited for this opportunity
- On pace for delivery of a rapidly constructed 2U
 - Start to finish 8 months!



ACKNOWLEDGEMENTS

- NASA ELaNa
- SRI International
- CXBN Team





COSMIC X-RAY BACKGROUND NANOSAT

