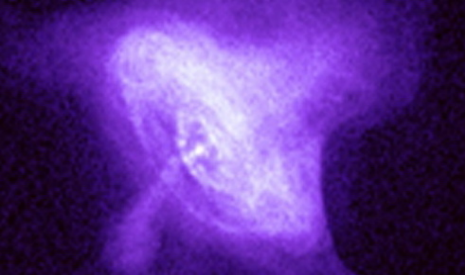


A Small Satellite Scientist Looks at Nanosat Capabilities

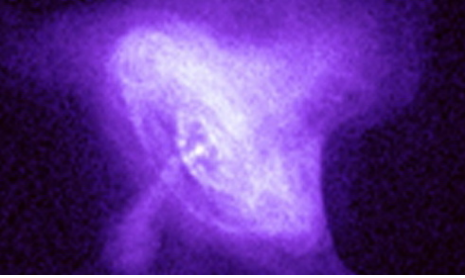
April 10, 2008

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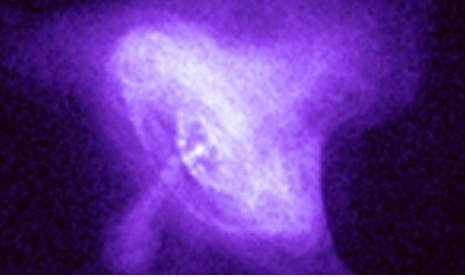
Galileo and Scaling

- Square-cube law (1683)
 - weight $\propto l^3$
 - strength $\propto l^2$
- Many other scalings favor small l
 - Thermal distortion
 - Speed and power of circuits
 - Telescope area/mass
 - Telescope area/\$



Scaling Drivers

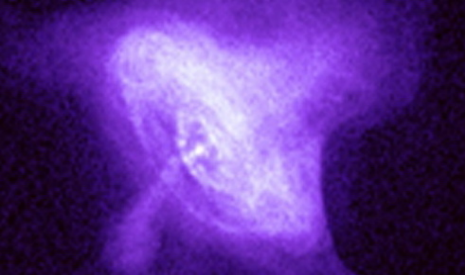
- Area
 - Available power
 - Heat flow
 - Telescope Speed
- Volume
 - Mass
 - Cost
 - Energy Storage



What Science (1)?

"Eighty percent of success is showing up." -- Woody Allen

- Gamma-ray bursts
 - 100 cm² detector enough for faint ones
 - Cosmological studies
 - 1 cm² detector enough for bright ones
 - Studies of the burst phenomenon
 - Major driver of detection rate is sky coverage, not detector area
- X-ray all sky monitors
 - RXTE ASM detectors only 60 cm²
 - Today's Si detectors better QE, resolution, allow shrinkage

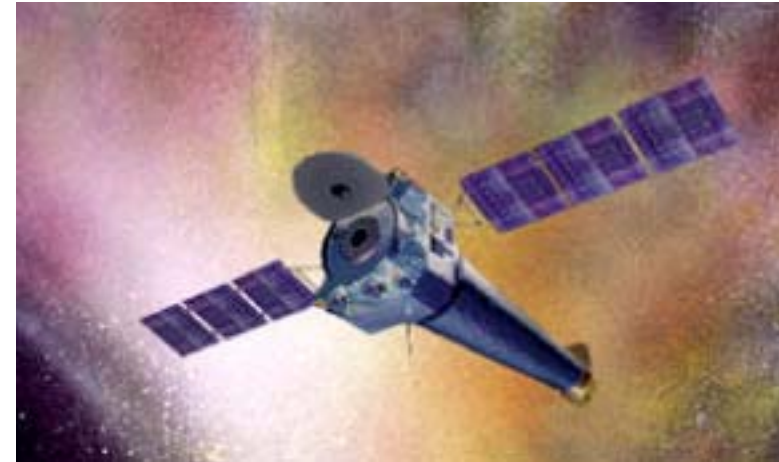


What Science (2)?

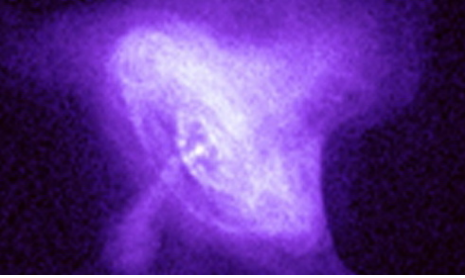
A spacecraft should do one thing well (to paraphrase Brian Kernighan)

- Chandra

- Cost ~\$4,000,000/cm²
- Excellent resolution
 - Heavy inefficient optics
 - Square-cube mechanical difficulties
- Many observations could exchange time for area
 - Nanosat could concentrate on a single target



What sets the floor?

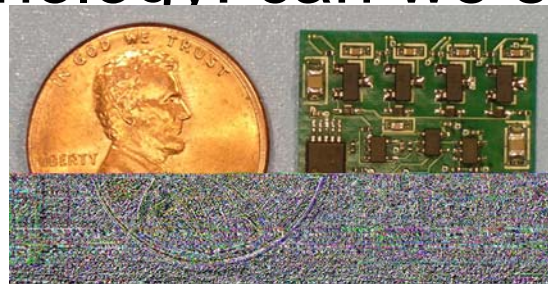


- Batteries

- Orbit time scale unchanged, energy needs scale as l^2 , energy capacity as l^3
 - NiCd adequate for HETE scale (125 kg)
 - Better batteries now available (nanophosphate Li look very promising)

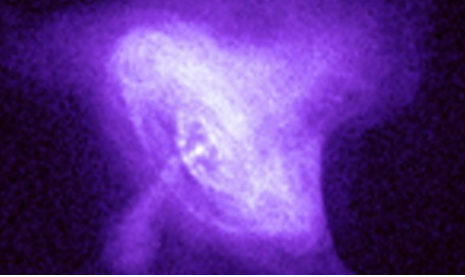
- Circuits

- HETE used 1992 technology: can we shrink area and power 25x? Yes:



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Conclusion

- In 1992, putting a world-class scientific mission into 125 kg was a challenge.
- In 2008, with the right choice of mission, 1 kg is the challenging scale.
- Keeping costs down is critical.
 - Still need to work out all details of a mission.
 - Need to publish detailed designs and software for reuse by the community.